

Efficient Computation of Global Illumination for Image Synthesis

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Abstract

We are exploring various functional approximation schemes (such as radial bases and Shepard’s interpolation) for precomputed radiance transfer (PRT), a finite element approach for real-time radiosity of complex scenes.

1. Introduction

Radiosity [1] is a general method for realistic rendering that uses finite element modeling to solve the Kajiya’s rendering equation [3]. In this formulation, the light flow in the scene is found by solving a large system of linear equations $(I - R)\lambda = \epsilon$, where the vector ϵ gives the spontaneous light emission and λ gives the total emission (spontaneous plus scattered) of each scene element.

Implementation of the radiosity method requires choosing a finite element scheme to model the light flow. The goal of our project is to compare various schemes, such as radial bases [2] and Shepard’s interpolation [5] basis - especially for use in real-time animation through precomputed radiance transfer [4, 6].

2. The rendering equation

The rendering equation can be written as $L = E + \mathcal{R}L$, where:

- $L(x, \vec{u})$ is the (unknown) radiance function, the total amount light emitted or scattered by the scene near the surface point x along directions near the unit vector \vec{u} ;
- $E(x, \vec{u})$ is the spontaneous emission function, the amount of light emitted by the scene near x and along \vec{u} ;
- \mathcal{R} is the light transfer operator, that represents the transport and interaction of the light with the scene’s objects. Its effect on an arbitrary function $F(x, \vec{u})$ is

$$(\mathcal{R}F)(x, \vec{u}) = \int_{\mathcal{S}} \rho(x, \vec{v}, \vec{u}) F(x', -\vec{v}) G(x, x', \vec{n}(x)) d\vec{v} \quad (1)$$

- \mathcal{S} is the set of all directions (i.e. the unit sphere).

- $x' = x^*(x, \vec{v})$ is the first point on the scene’s surface found by a ray that leaves x towards \vec{v} .

- $\rho(x, \vec{v}, \vec{u})$ is the scattering coefficient, the fraction of the incident light at the point x , coming from the direction \vec{v} , that is re-emitted along directions near \vec{u} ;

- $G(x, x', \vec{n}(x))$ is the geometric factor that depends on the orientation of the surface at x .

In favorable circumstances, the solution of the rendering equation is

$L = (\mathcal{I} - \mathcal{R})^{-1}E$, where \mathcal{I} is the identity operator. The rendering operator $(\mathcal{I} - \mathcal{R})^{-1}$ can be computed by Neumann’s formula

$$(\mathcal{I} - \mathcal{R})^{-1} = (\mathcal{I} + \mathcal{R} + \mathcal{R}^2 + \mathcal{R}^3 + \dots) \quad (2)$$

Each term \mathcal{R}^k accounts for light that interacted k times with the surface of the scene before being observed.

3. Finite element radiosity

We use the term *site* to mean a pair $p = (\hat{p}, \vec{p})$ where \hat{p} is a point on the scene’s surface and \vec{p} is the corresponding unit normal vector. For Lambertian radiosity, a finite element is a function ϕ defined on the scene’s surface sites, such that $\phi(p)$ is nonzero only for a relatively small and compact set of sites (the support of ϕ).

In point-based finite element methods, the basis ϕ is defined by selecting a number of sampling sites $P = \{p_1, p_2, \dots, p_n\}$ on the scene’s surface (see Figure 1(a)) and then choosing for each p_i a corresponding finite element ϕ_i that has p_i as its centroid (see Figure 1(b)).

3.1 Basis functions

In a radial basis each element ϕ_i is derived from fixed mother function ϕ by the formula $\phi_i(p) = \Phi(\|p, p_i\|/\alpha_i)$; where $\|p, q\|$, denotes the distance between the sites p and q , and α_i is a parameter that determines the “mean radius” of the element. See figure 2(a).

In Shepard’s basis [5], each element is

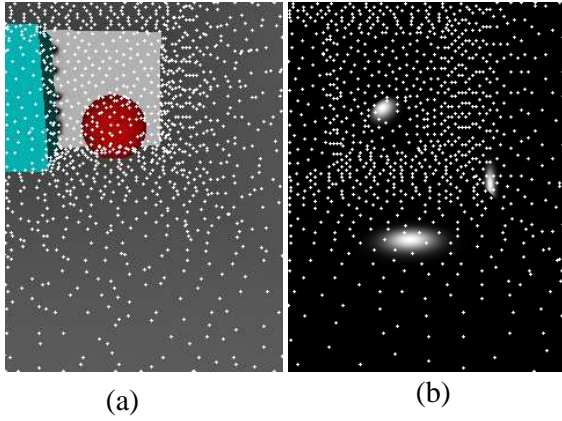


Figure 1. The point-based FE method.

$$\phi_i(p) = \frac{w(p, p_i)}{\sum_{j=1}^n w(p, p_j)}$$

Here $w(p, q)$ is a non-negative function that tends to infinity when p tends to q . e. g. $1/\|p, q\|$. See figure 2(b).

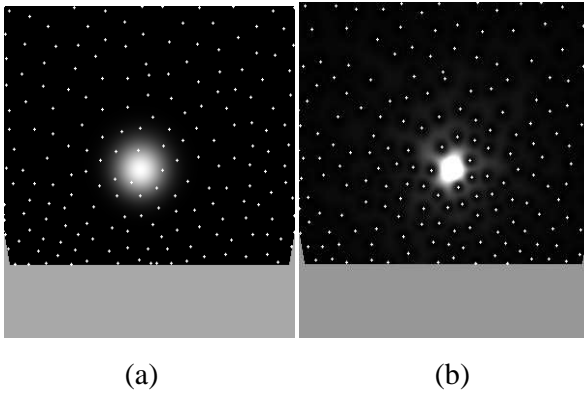


Figure 2. Radial and Shepard elements.

4. Discretizing the rendering equation

In finite element radiosity, the rendering equation becomes a linear equation system $\lambda = \epsilon + R\lambda$, where $\lambda = (\lambda_1, \lambda_2, \dots, \lambda_n)$ are the coefficients of L in the chosen basis, and $\epsilon = (\epsilon_1, \epsilon_2, \dots, \epsilon_n)$ are the coefficients of E . Each element R_{jk} represents the influence of the light emitted by ϕ_k on the light scattered by ϕ_j . It is determined by the formula

$$R_{jk} = G(\hat{p}_k, \hat{p}_j, \vec{p}_j) \rho(p_j) V(\hat{p}_k, \hat{p}_j) \quad (3)$$

Once the matrix R is available, the coefficients λ_i of the radiance function L can be computed by setting $\lambda \leftarrow (0, \dots, 0)$ and then iterating $\lambda \leftarrow \epsilon + R\lambda$ until convergence.

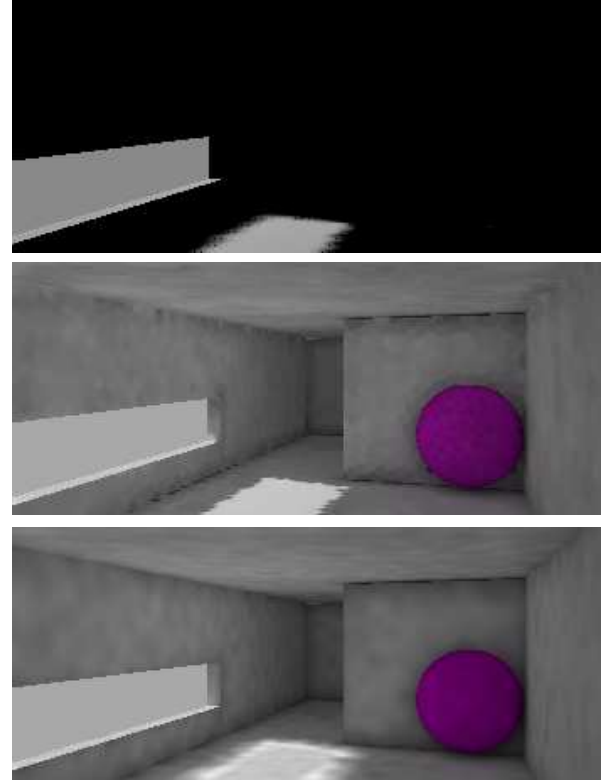


Figure 3. Rendered scenes.

5. Results

Figure 3 shows a test scene rendered without radiosity (top) and with ten radiosity iterations, using Shepard interpolation (middle) and a radial basis (bottom).

References

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